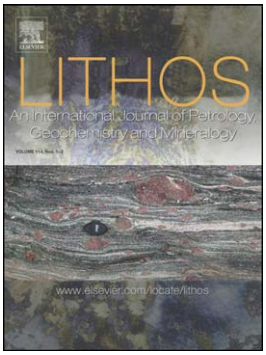

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Geochronology and geochemistry of the Early Jurassic Yeba Formation volcanic rocks in southern Tibet: initiation of back-arc rifting and crustal accretion in the southern Lhasa Terrane

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Abstract

Understanding the geological history of the Lhasa Terrane prior to the India-Asia collision (~55±10 Ma) is essential for improved models of syn-collisional and post-collisional processes in the southern Lhasa Terrane. The Miocene (~18–10 Ma) adakitic magmatism with economically significant porphyry-type mineralization has been interpreted as resulting from partial melting of the Jurassic juvenile crust, but how this juvenile crust was accreted remains poorly known. For this reason, we carried out a detailed study on the volcanic rocks of the Yeba Formation (YF) with the results offering insights into the ways in which the juvenile crust may be accreted in the southern Lhasa Terrane in the Jurassic. The YF volcanic rocks are compositionally bimodal, comprising basalt/basaltic andesite and dacite/rhyolite dated at 183-174 Ma. All these rocks have an arc-like signature with enriched large ion lithophile elements (LILEs; e.g., Rb, Ba and U) and light rare earth elements (LREEs) and depleted high field strength elements (HFSEs; e.g., Nb, Ta, Ti). They also have depleted whole-rock Sr-Nd and zircon Hf isotopic compositions, pointing to significant mantle isotopic contributions. Modeling results of trace elements and isotopes are most consistent with the basalts being derived from a mantle source metasomatized by varying enrichment of subduction components. The silicic volcanic rocks show the characteristics of transitional I-S type granites, and are best interpreted as resulting from re-melting of a mixed source of juvenile amphibole-rich lower crust with

reworked crustal materials resembling metagraywackes. Importantly, our results indicate northward Neo-Tethyan seafloor subduction beneath the Lhasa Terrane with the YF volcanism being caused by the initiation of back-arc rifting. The back-arc setting is a site for juvenile crustal accretion in the southern Lhasa Terrane.

Keywords: geochemistry; geochronology; Sr-Nd-Hf isotopes; back-arc; Yeba Formation; crustal accretion; Tibet

1 Introduction

The Tibetan Plateau records a complete set of tectonic scenarios from continental rifting to microcontinent drifting, to continent-continent collision, to continental amalgamation and to the ultimate uplift, formatting the largest plateau on the Earth (Allegre et al., 1984; Dewey et al., 1988; Yin and Harrison, 2000; Zhu et al., 2011a, b, 2013). The Lhasa Terrane, which serves as the southern margin of the Asian continent before the India-Asia collision, is an important constituent of the Tibetan Plateau. Recent studies of the Miocene adakitic magmatism with significant porphyry-type mineralization (Hou et al., 2015; Tang et al., 2010) require the potential presence of Jurassic juvenile crust in the Lhasa Terrane (Hou et al., 2015; Zhu et al., 2011a, b, 2013). Where such juvenile crustal rocks resided and how they were emplaced become important questions. In the modern Earth, crustal growth is thought to primarily take place along convergent plate boundaries where granitoid magmas result from advanced extent of fractional crystallization of basaltic magmas or mostly from re-melting of mantle-derived basaltic rocks (Moyen et al., 2016; Niu et al., 2013 and references therein). The southern Lhasa Terrane has been considered as a prolonged active continental margin in the Mesozoic, which would be the site for crustal growth as per the popular view. However, among other reasons, active subduction zones such as island and continental arcs (i.e., active continental margins) are not sites for net crustal accretion because crustal production and destruction (sediment recycling and

subduction erosion) are mass balanced (Niu et al., 2013). The key condition for net crustal growth is the preservation of juvenile crust, which is possible in broad settings of continental collision (Niu et al., 2013).

Existing studies all suggest that the magmatism of early Cretaceous to early Tertiary age in the southern Lhasa Terrane resulted from northward subduction of the Neo-Tethyan seafloor (including effects of slab rollback and breakoff; Chu et al., 2006; Guo et al., 2014; Ji et al., 2009; Kang et al., 2014; Wang et al., 2016; Zhu et al., 2008, 2009). However, the nature and tectonic setting of the Triassic – late Jurassic magmatism has been an issue of debate. Two models have been proposed to explain the magmatism of this earlier time. One view is that the northward Neo-Tethyan seafloor subduction had already commenced since the late Triassic and can account for this earlier stage of magmatism (Chu et al., 2006; Guo et al., 2014; Kang et al., 2014; Wang et al., 2016; Zhu et al., 2008). The other view is that this earlier magmatism may have resulted from southward subduction of the Bangong-Nujiang Tethyan seafloor (Zhu et al., 2011b, 2013).

In this study, we offer solutions to the above problems on the basis of our comprehensive study on the bimodal volcanic rocks from the Jurassic Yeba Formation in the southern Lhasa Terrane. These include (1) zircon U-Pb age data, (2) zircon in-situ Hf isotopes, (3) bulk-rock major and trace element compositions, (4) bulk-rock Sr-Nd isotopes, and (5) logical interpretations of these data.

2 Geological background and sample details

The Lhasa Terrane is the southernmost tectonic unit of the Asian continent, bounded by the Bangong-Nujiang Suture Zone (BNS) to the north and the Indus-Yarlung-Zangbo Suture Zone (IYZSZ) to the south (*Fig. 1a, b*), and is considered a fragment of Gondwana that drifted over the Tethyan Ocean basin and accreted to the Asia continent long before the India-Asia collision (Pan et al., 2012; Sengör, 1987; Yin and Harrison, 2000; Zhu et al., 2011a, b, 2013). The Lhasa Terrane comprises the northern, central, and southern subterranes, separated by the Shiquan River-Nam Tso mélange zone (SNMZ) and the Luobadui-Milashan fault (LMF), respectively (Zhu et al., 2011b). The three subterranes have different sedimentary covers and metamorphic basements. The central Lhasa subterrane is underlain by a Proterozoic-Early Cambrian crystalline basement represented by the Nyainqentanglha Group, and is covered with Permian-Carboniferous metasedimentary rocks and late Jurassic – early Cretaceous volcano-sedimentary sequences (cf. Pan et al., 2012; Wang et al., 2007). The southern and northern Lhasa subterranes have different crustal structures that are dominated by subduction-related juvenile basements with rare ancient components (Hou et al., 2015; Zhu et al., 2011b, 2013). The most remarkable magmatism in the Lhasa Terrane is illustrated by the Gangdese magmatic belt, which extend E-W for ~1500 km in the southern Lhasa subterrane and is composed of the late Triassic – early Tertiary Gangdese batholiths and the Mesozoic-Cenozoic volcanic successions (cf. Chu et al., 2006; Hou et al., 2015; Mo et al., 2003; Ji et al.,

2009; Zhu et al., 2011a, 2013).

The Mesozoic volcanic rocks in the Gangdese magmatic belt include the early Jurassic – early Cretaceous Sangri Group (SG) and the early Jurassic Yeba Formation. The SG, outcropped to the north of the IYZSZ, is dominated by basalts, basaltic andesites, andesites, minor dacites and rhyolites with intercalations of sandstone, slate and limestone (cf. Kang et al., 2014; Zhu et al., 2009). Zircon U-Pb dating gives the Sangri Group volcanic rocks a wide age range of 195-137Ma (Kang et al., 2014; Zhu et al., 2009). Recently, Wang et al., (2016) identified middle-late Triassic volcanic rocks near the Changguo village with zircon U-Pb ages of 237-211Ma, which are previously assigned to the Sangri Group. Apart from the volcanic suites mentioned above, a special block named "Zedong terrane" (ZT) was identified by Aitchison et al., (2000). The ZT exposed only in the Zedong area (~25 km^2), which lies in fault contact with the ISZYZ in the south and with the Cenozoic Luobusha conglomerate in the north (cf. Hu et al., 2016). It is composed of basalts, basaltic andesites, andesites, rare dacites with other intrusives (e.g. gabbro, hornblendite, granitoids; Aitchison et al., 2000; Zhang et al., 2014). Geochronological studies show that this terrane was formed at the Late Jurassic (160-155Ma; McDermid et al., 2002; Zhang et al., 2014). Recent studies suggest the ZT may preferably represent a slice of magmatic arc developed on the southern margin of the Lhasa Terrane (Hu et al., 2016; Zhang et al., 2014) rather than an intra-ocean island arc within the Neo-Tethyan ocean (Aitchison et al., 2000, 2007).

The Yeba Formation (YF) crops out north of the SG, extending E-W for ~250 km from Dagze to Gongbo-Gyamda County. The YF is disconformably overlain by the Upper-Jurassic Duodigou Formation and Cretaceous Menzhong Formation, and is intruded locally by plutons of the Gangdese batholiths. The YF volcanic sequence is compositionally bimodal, dominated by basalts (~3000 m thick) and silicic lavas (~2000 - 7000 m thick) with rare intermediate members. The sedimentary layers mainly occur in the uppermost section of the YF, comprising fine-grained sandstone, calcic slate, bioclast limestone interbedded with siliceous rocks (cf. Pearce et al., 1988; Zhu et al., 2008). Most of the YF volcanic rocks have undergone up to greenschist-facies metamorphism. The zircon U-Pb dating gives the YF volcanic rocks the early Jurassic age (192-174Ma; cf. Dong et al., 2006; Chen et al., 2009; Zhu et al., 2008).

Twenty-two representative volcanic samples from the YF were collected for study. Sample locations and a representative outcrop are shown on *Fig.1c, d*. Most of the basaltic rocks have a porphyritic texture with phenocrysts of plagioclase ± olivine ± clinopyroxene whereas the silicic rocks have a rhyotaxitic texture (*Fig. 1e-g*). A common secondary mineral assemblage of greenschist facies overprint includes chlorite ± epidote ± calcite ± sericite ± magnetite in all the samples. Basalts from the Baiding area contain peridotite xenoliths (*Fig. 1f*).

3 Analytical methods

3.1 Zircon U-Pb age dating and Hf isotope

Zircons were extracted using standard density and magnetic methods. Separated zircon grains were handpicked under a binocular microscope before mounted in epoxy resin and polished. Cathodoluminescence (CL) images were taken to study their internal structures and for better target site selection.

Zircon age dating was analyzed using Laser ablation inductively coupled plasma mass spectrometry (LA-ICP-MS) at the State Key Laboratory of Geological Processes and Mineral Resources, China University of Geosciences, Wuhan (CUGW). An Agilent 7500 ICP-MS coupled with a laser ablation system (GeoLas 2005) was used for zircon U-Pb isotope and trace element analysis. Zircon grains were ablated by using a spot diameter of 32μm with acquisition time of 20s for background and 50s for data collection. The ablated aerosol was carried by high-purity helium gas and introduced into the ICP-MS. Zircon standard 91500 was used as an external standard to monitor isotopic fractionation during analysis. The $^{207}Pb/^{206}Pb$ and $^{206}Pb/^{238}U$ ratios were calculated offline using *ICPMSDataCal_ver9.0* (Liu et al., 2010) with the weighted mean U-Pb ages calculated, and the Concordia plots constructed, by using *Isoplot_ver4.15* following Liu et al. (2008a). The Hf isotope ratio analysis was carried out on a Neptune multi-collector (MC)-ICP-MS coupled with a 193nm laser (GeoLas Plus) at the Institute of Geology and Geophysics, Chinese Academy of Sciences

(IGGCAS) in Beijing following Wu et al. (2010). The $^{176}Hf/^{177}Hf$ ratios of standard zircon 91500 gave a weighted mean value of 0.282292±6 (2σ) consistent with the recommended value within error. The U-Pb age data and Hf isotopic analyses are given in the online appendices *Table S1* and *Table S2*, respectively, with reference materials (91500 and GJ-1) being plotted in *Figure S1*.

3.2 Whole rock major and trace element geochemistry

A total of 22 samples were selected for major and trace element analysis. Major and trace element concentrations were measured using X-ray fluorescence (XRF) spectroscopy and inductively coupled plasma mass spectrometry (ICP-MS), respectively at CUGW. The sample powders were dried at 105℃ for 12 hours. Mixtures of 0.5g rock powder and 5g $LiBO_2$ + $Li_2B_4O_7$ + NH_4NO_3 + LiBr had been heated at 1000°C in a Pt crucible to make fused glass disks, and were later analyzed on XRF-1800. Loss on ignition (LOI) was determined after heating ~ 1g rock powder in a Muffle furnace at 1000°C for 2 hours. See Ma et al. (2012) for analytical details. For trace elements analysis, dried sample powders of 50 ± 1mg were dissolved in Teflon bombs using high-purity HNO_3 + HF and heated at 190 ± 5°C for over 48h. The chilled solutions were then dried by distillation at 140°C before the residual was re-dissolved using 1ml HNO_3 and re-dried again. The resultant residues were again dissolved using 3ml 30% HNO_3, heated at 190 ± 5°C for more than 12h before transferred to polyethylene bottle and diluted to 100g 2% HNO_3 solutions for analysis

on Agilent 7500a. The standards AGV-2, BHVO-2, BCR-2 and RGM-2 were used to monitor data quality during the analysis. Analytical details and data precision/accuracy are given in Liu et al. (2008b). The major and trace element data are given in *Table 1*. Analyses of procedural blank, reference materials and replicates are given in *Table S3* and *Table S4*.

3.3 Whole rock Sr-Nd isotopes

The Sr-Nd isotopic analysis was performed on 19 samples. Separation and purification of Sr and Nd were carried out using conventional ion exchange techniques in the Key Laboratory of Orogenic Belts and Crustal Evolution at Peking University. The Sr and Nd isotopic analysis was performed using thermal ionization mass spectrometry (TIMS; Thermo-Finnigan TRITON) in Tianjin Institute of Geology and Mineral Resources following the procedure by Liu et al. (2017). The Sr isotopic ratios were normalized to $^{86}Sr/^{88}Sr=0.1194$ whereas the Nd isotopic ratios were normalized to $^{146}Nd/^{144}Nd=0.7219$ for mass fractionation correction. Repeated analyses of reference standard BCR-2 and NBS987 give weighted mean $^{87}Sr/^{86}Sr$ = 0.704978±38 (2σ, n=7) and 0.710218±33 (2σ, n=8), respectively (*Fig. S2*). The weighted mean $^{143}Nd/^{144}Nd$ values of BCR-2 and a laboratory internal standard solution (LRIG-Nd) yielded 0.512635±2 (2σ, n=7) and 0.512201±2 (2σ, n=8), respectively, which are the same as recommended values within error (Li et al., 2007). The analytical results are given in *Table 1*.

4 Results

4.1 Zircon U-Pb geochronology and Hf isotope

Zircons from one basaltic and five silicic volcanic samples were selected for U-Pb dating. Most zircon grains are euhedral with oscillatory zoning, and have varying size (~ 25 - 140 μm) and elongation ratio (~2 – 5), reflecting their magmatic origin (Corfu et al., 2003). Zircons from the basaltic sample (YB1323, *Fig. 4a*) give a weighted mean $^{206}Pb/^{238}U$ age of 178.0±1.3Ma (2σ). Zircons from four dacite samples (YB1310, *Fig. 4b*; YB1320, *Fig 4d*; YB1324, *Fig. 4e*; YB1326, *Fig. 4f*) and one rhyolite sample (YB1318, *Fig. 4c*) give weighted mean ages of 183-174Ma.

A subset of 85 Hf isotopic analyses of zircon grains are plotted in *Fig. 5a*. The initial Hf ratios are calculated using their corresponding U-Pb ages for all the analytical spots. Zircon grains from the basaltic andesite sample (YB1323) show initial $^{176}Hf/^{177}Hf$ ratios of 0.282684-0.283000 ($\varepsilon_{Hf}(t)$=0.8-11.9), and those from four silicic samples (YB1310, YB1318, YB1320, YB1324) yield similar values (initial $^{176}Hf/^{177}Hf$=0.282728-0.283028, $\varepsilon_{Hf}(t)$=2.4-13.0). One silicic sample displays more depleted Hf isotopes ($\varepsilon_{Hf}(t)$=11.2-17.6), suggesting a heterogeneous source for the silicic magmas.

4.2 Whole-rock major and trace element geochemistry

The 22 samples show large SiO_2 (on an anhydrous basis) variation (~ 43.5 - 74.0

wt.%) with a gap in the intermediate compositions (~ 56.0 - 65.5 wt.%). Because most samples have undergone varying extents of alteration, the Nb/Y-Zr/TiO_2 diagram (Winchester and Floyd, 1977) is used as proxy for the total alkali-silica (TAS) diagram (Le Bas et al., 1986) and the Co-Th diagram (Hastie et al., 2007) as proxy for the SiO_2-K_2O diagram (Peccerillo and Taylor, 1976). The samples plot in two separate groups: a basaltic group composed of calc-alkaline basalt/basaltic andesite, and a silicic group with dacite/rhyolite compositions (*Fig. 2*).

4.2.1 Basaltic group

The basaltic samples have varying SiO_2 (43.5-52.8 wt.%), MgO (5.02-10.65 wt.%), Cr (44.9-223 ppm) and Ni (29.2-91.4 ppm), suggesting their variably evolved nature. Sample YB1302 and YB1303 have very low SiO_2 (43.5-43.9%), probably caused by post-magmatic alteration. Sample YB1319 and YB1323 have higher SiO_2 plotting in the basaltic andesite field (*Fig. 2b*). All the samples are TiO_2 poor (0.88-1.44 wt.%) with relatively high Al_2O_3 (14.89-18.06 wt.%). Calcite and chlorite are ubiquitous secondary minerals that are probably responsible for their high LOI (loss on ignition) values. An N-MORB normalized multi-element diagram and chondrite normalized REE patterns are shown in *Fig.3*. The Yeba basaltic samples are enriched in large ion lithophile elements (LILEs) and light rare earth elements (LREEs), but depleted in high field strength elements (HFSEs) with no obvious Eu anomaly (Eu/Eu* = 0.89-1.05). That is, the basalts have an arc-like geochemical signature (Pearce and

Peate, 1995; Perfit et al., 1980).

4.2.2 Silicic group

The twelve silicic samples are compositionally dacitic/rhyolitic (*Fig. 2*) and metaluminous to peraluminous (aluminum saturation index, ASI=0.94-1.67) with low MgO and varying alkalis (Na_2O/K_2O=0.2-2.89, Na_2O+K_2O=3.36-6.65 wt.%). A well-defined correlation between K_2O and Na_2O (not shown) suggest that the alkalis were largely unaffected by weathering process. The silicic samples are enriched in LILEs and depleted in some HFSEs (i.e., Nb, Ta and Ti vs. Zr and Hf; *Fig. 3b*) with weak REE fractionation and weak negligible Eu anomalies (Eu/Eu* = 0.71-1.04; *Fig. 3d*).

4.3 Whole rock Sr-Nd isotopes

The Sr-Nd isotopes are plotted in *Fig. 5b*. The basaltic rocks show homogeneous Sr isotopic compositions ($^{87}Sr/^{86}Sr(t)$=0.7039-0.7049, t=174Ma), while the silicic rocks show a wider range ($^{87}Sr/^{86}Sr(t)$=0.7027-0.7053). The Nd isotopes also show minor variations in both groups ($\varepsilon_{Nd}(t)$=1.6-4.0 for the basaltic group and $\varepsilon_{Nd}(t)$=0.9-3.4 for the silicic group).

5 Discussion

5.1 Petrogenesis of the YF basaltic rocks

5.1.1 Testing trace element mobility and crustal contamination

Since the YF basaltic rocks have been variably altered, it is necessary to evaluate the effect of alteration on trace element mobility to prevent from misinterpretation of the data. In this study, Zr is chosen as an immobile element to test the mobility of other elements (Cann, 1970; Hastie et al., 2007; Hill and Worden, 2000). Other HFSEs and REEs exhibit immobile nature because of their good correlations with Zr (not shown) and subparallel patterns in both REE and multi-element diagrams (*Fig. 3a, c*). However, LILEs in the basaltic samples (e.g., Rb, Ba) are obviously mobile as shown by their weak correlations with Zr (not shown) and random variability (*Fig. 3c*). Therefore, except for LILEs, most other trace elements can be used to evaluate the magma generation and differentiation processes.

Some of the basaltic samples are quite evolved, especially for basaltic andesite samples YB1319 and YB1323 with elevated SiO_2, low MgO, Cr and Ni, which might be caused by AFC (assimilation - fractional crystallization; DePaolo, 1981). However, it is unlikely that AFC is responsible for the compositional variations of the YF basalts because (1) the basalts have much lower abundances of “crustal-affinity” elements (e.g. Th=1.58 ppm, U=0.55 ppm, K=0.53 wt.% on average) than those of average upper crust (cf. Rudnick and Gao, 2003), which suggests a limited addition of

crustal materials during post melting processes (Zhu et al., 2008); (2) The essentially constant $^{87}Sr/^{86}Sr(t)$ with varying MgO suggests a tendency towards fractional crystallization without crustal contamination as otherwise an inverse correlation would be expected (*Fig. 6a)*; (3) The lack of correlation between $\varepsilon_{Nd}(t)$ and Th/Nb (*Fig. 6b*) suggests the addition of subduction components instead of crustal materials with low $\varepsilon_{Nd}(t)$ (~ -10.2; Wen et al., 2008). It follows that mantle source heterogeneity and subduction components influx are likely responsible for the compositional variations of the basalts.

5.1.2 The nature of mantle source

The enrichment of LILEs relative to HFSEs is consistent with the source being mantle wedge with added slab components (Elliott, 2003; McCulloch and Gamble, 1991; Pearce and Peate, 1995). The slab components include (1) aqueous fluids from hydrous oceanic crust and/or overlying sediments; and (2) silicate melts from oceanic basalt and/or sediments (Hastie et al., 2010; Pearce and Peate, 1995; Perfit et al., 1980). It is consensus that aqueous fluids dehydrated from the slab can readily concentrate and transport LILEs, leaving HFSEs in the slab because of the immobility of the latter, whereas silicate melts can carry most of the incompatible elements (Keppler, 1996; Tatsumi et al., 1986). A lithologic association of high-Mg andesite – adakitic TTG/D (tonalite-trondhjemite-granodiorite/dacite) is considered as resulting from reaction between slab-derived silicate melts and mantle peridotite in the arc

setting (Drummond et al., 1996; Polat and Kerrich, 2011; Wang et al., 2007). In this study, the absence of this rock association in the YF suggests that slab-derived silicate melts were not involved in the magma generation (Hastie et al., 2010). Actually, such a lithologic association exists in the adjacent Zedong terrane (ZT) and the Sangri Group (SG) (Zhang et al., 2014; Zhu et al., 2009), which indicates the involvement of slab-derived melts in their genesis. Therefore, we suggest that the HFSEs in the YF basalts were largely inherited from the mantle sources with negligible exotic addition.

The YF basalts have MORB-like HFSEs and HREEs abundances (*Fig. 3c*), reflecting melting in the spinel peridotite facies. Figure *6f* shows partial melting curves of spinel lherzolite (with source mode of Ol:Opx:Cpx of 54:30:16 and melting mode of 0:30:70; Gribble et al., 1998) with primitive mantle (PM, Sun and McDonough, 1989) and depleted MORB mantle (DMM, Salters and Stracke, 2004) compositions. The curves are calculated using non-modal batch melting model (Shaw, 1970) with partitioning coefficients from Pearce and Parkinson (1993). An interval of 5-15% partial melting is required to form these basalts. We suggest that such degree of partial melting only has a small effect on the incompatible element ratios, as reflected by the essentially constant La/Sm with varying La (*Fig. 6e*). Therefore, ratios of incompatible elements can be used to illustrate the nature of the mantle source. The YF basalts are plotted in the diagrams of Nb vs. Nb/Ta and Zr/Yb vs. Nb/Ta (*Fig 6c, d*), with adjacent arc-like basalts from the SG (Sangri Group, Kang et al., 2014) and the ZT (Zedong terrane, Zhang et al., 2014) plotted for comparison.

Although Nb and Ta are thought to be geochemically similar, it has been observed that Nb is incompatible than Ta in seafloor basalts and peridotites (see review by Niu, 2012) as also confirmed experimentally (Forsythe et al., 1994; Green, 1994). This means that the low Nb/Ta ratios (usually subchondritic) at low Nb abundances in arc basalts indicate a prior mantle source depletion in a melt component with superchondritic Nb/Ta ratios (Eggins et al., 1997; Elliott, 2003; Leat et al., 2003). Furthermore, these preferential melt-extraction episodes have been manifested in many back-arc regions associated with individual arcs (Leat et al., 2003; McCulloch and Gamble, 1991; Pearce and Stern, 2006; Pearce et al., 2005; Woodhead et al., 1993). It is likely the case in this study because of the following reasons. (1) The YF basalts are characterized by superchondritic Nb/Ta ratios (~18.5-21.0), whereas those from the adjacent arcs (represented by the SG and ZT volcanic rocks) are subchondritic (~14.6-17.2 and ~7.27-13.2, respectively) with relatively lower Nb abundances (*Fig. 6c*). (2) The gradient shown in the Zr/Yb vs. Nb/Ta diagram (*Fig. 6d*) indicates that the mantle source of the Yeba basalts was enriched in HFSEs relative to those in the adjacent arcs, as Zr/Yb ratio is used as a proxy for melt extraction (Leat et al., 2003; Pearce et al., 1995). The mantle source of the adjacent arcs became depleted due to the preferential loss of small melt fractions in the back-arc region, where the Yeba basaltic magmas had been generated (see below). (3) The bimodal feature of the YF volcanic suite suggests an extensional tectonic setting, whereas the SR and ZT have a large volume of andesites exhibiting typical island-arc lithologic association

(Kang et al., 2014; Wang et al., 2016; Zhang et al., 2014). It is noteworthy that similar lithologic association was identified in the Ryukyu arc – Okinawa back-arc basin system (e.g., Shinjo et al., 2000). Moreover, the YF basalts compositionally resemble those from the Okinawa (*Fig. 3a, c*) and Mariana troughs, which have been demonstrated to form in the early stages of back-arc basin development (Gribble et al., 1998; Shinjo et al., 2000).

We thus suggest that the YF volcanism be best interpreted as resulting from initiation of a back-arc basin in the southern Lhasa Terrane.

5.2 Origin of the YF silicic rocks

It is widely accepted that silicic magmas can be generated in arc-related settings along the following pathways: (1) fractional crystallization from associated arc basaltic magmas, probably accompanied by crustal assimilation (DePaolo, 1981; Haase et al., 2006; McCulloch et al., 1994; Shinjo and Kato, 2000); (2) partial melting of mid/lower crust that has composition equivalent to meta-basalts to meta-dacites (Patiño Douce and Beard, 1995; Shukuno et al., 2006; Smith et al., 2003; Tamura and Tatsumi, 2002). Either of the two potential processes could result in similar major and trace element compositions of silicic magmas with isotopic similarities if reworked crustal components were absent. The YF silicic rocks have similar Sr-Nd isotopic characteristics as those of the basalts. Thus, it is not straightforward to distinguish the two processes. However, we suggest that the silicic rocks were more likely generated

through partial melting rather than AFC from the associated basaltic rocks because (1) there is no compositional continuity expected for fractional crystallization between the basaltic and silicic rocks on harker diagrams (not shown); (2) the scatter is greater than can be explained by fractional crystallization in the La vs. La/Sm diagram (*Fig. 6e*); (3) the silicic volcanic rocks are widespread and volumetrically significant, which cannot be explained by fractional crystallization from associated arc basaltic magmas (Zhu et al., 2008).

It is consensus that the compositional diversity of silicic melts depends on varying source compositions and melting conditions. For example, it is widely accepted that partial melting of metasedimentary sources forms S-type granites whereas I-type granites originate from metaigneous sources (Chappell and White, 2001); and that melts derived from thickened lower crust or subducted oceanic crust usually have adakitic characteristics if garnet is present as a residual phase (Defant and Drummond, 1990; Drummond et al., 1996; Hou et al., 2004). The concave-upward REE patterns for the silicic rocks (*Fig. 3b*) together with the high Al_2O_3 and negligible Sr, Ba and Eu anomalies imply that plagioclase would be largely melted out from the source region leaving amphibole-pyroxene rich residues rather than garnet-bearing residues. This indicates a relatively low pressure condition during magma generation and segregation (Altherr et al., 2000; Patiño Douce, 1995; Patiño Douce, 1999). In terms of isotopic evidence, Zhu et al. (2008) proposed that the silicic rocks originated from a hybridized source that was composed of juvenile amphibole-bearing lower crust

mixed with a small amount of reworked crustal material. The zircon Hf isotopic features also reveal the silicic rocks to be dominated by depleted Hf isotopic compositions with a mixing tendency towards reworked components (*Fig. 5a*).

Experimental studies demonstrated that partial melts derived from different protoliths under variable melting conditions can be identified on the basis of major element oxides (Altherr et al., 2000; Patiño Douce, 1995, 1999; Patiño Douce and Beard, 1995). Figure *7c* shows that the silicic rocks might be derived from meta-mafic rock source. However, it is suggested based on experimental studies that dehydration melts from amphibolites (equivalent to meta-basalts) are compositionally similar to trondhjemitic rocks (Barker and Arth, 1976; Johannes and Holtz, 1996), which are distinct from the YF silicic rocks in Na_2O content (*Fig. 7a, b*). Moreover, melts derived from quartz amphibolites (compositionally similar to meta-andesite to meta-dacite) are often peraluminous (Patiño Douce and Beard, 1995), which are compositionally different from those of the YF silicic rocks. The relatively high $CaO/(MgO+FeO_T)$, low K_2O/Na_2O and ASI of the silicic rocks rule out the affinity to a metapelitic (MP) source (Altherr et al., 2000), but show more compatibility with melts of a metagraywacke (MGW) source (*Fig. 7b*). These features are analogous to those of the Cordilleran calc-alkaline granites as resulting from re-melting of a meta-basaltic source that had been hybridized with immature meta-sediments (i.e. metagraywackes) at convergent continental margins (Patiño Douce, 1999). Figure *7d* shows that data of the YF silicic rocks fit well with the curve that is constructed by the

reaction of meta-basalts and metagreywackes at low pressure (≤5kbar). Considering these data and the zircon saturation temperatures of 762-855°C calculated using Watson and Harrison (1983), it is suggested that the partial melts were in equilibrium with amphibole before segregation, which is consistent with the concave-upward REE patterns of the YF silicic rocks (*Fig. 3b*). Therefore, we propose that the YF silicic rocks have the characteristics of transitional I-S type granites, and originated from a hybridized source of meta-basalts and metagreywackes; the latter probably represent the "reworked crustal components" in the source of the silicic melts.

5.3 Jurassic Subduction polarity in the southern Lhasa Terrane

Two tectonic models have been proposed to explain the Mesozoic magmatism in the southern Lhasa subterrane, in which two reversed subduction polarities are postulated. One model assumes that the southern Lhasa Terrane was developed at an Andean-type orogen as a result of northward subduction of the Neo-Tethyan seafloor (Chu et al., 2006; Guo et al., 2014; Ji et al., 2009; Kang et al., 2014; Wang et al., 2016; Zhu et al., 2008). The other model hypothesizes that the southern Lhasa Terrane represents a passive margin of the Neo-Tethyan back-arc basin rather than an active continental margin during the late Triassic, where the magmatism was interpreted as resulting from re-melting of juvenile crust in a back-arc extensional region caused by southward subduction of the Bangong-Nujiang Tethyan seafloor (Zhu et al., 2011b, 2013). On the basis of our petrological and geochemical study presented above, we

offer new insights into the deep processes beneath the Lhasa Terrane and propose explicitly that northward subduction of the Neo-Tethyan seafloor must have been responsible for the Jurassic-Cretaceous magmatism in the southern Lhasa Terrane, for which we elaborated as follows.

(1) Direction of mantle flow. The asthenospheric mantle lost its fusible materials episodically as small melt fractions while flowing towards the trench, which was reflected in the gradient of Nb/Ta ratios (*Fig. 6g*) for the basalts. The refractory materials were subsequently transported to source region of the volcanic front by mantle convection (Eggins et al., 1997; Elliott, 2003; Leat et al., 2003). This suggests that the mantle flowed southward to trench side beneath the Lhasa Terrane, corresponding to northward subduction of the Neo-Tethyan Ocean slab as illustrated in *Fig. 8*.

(2) Influx of slab components. Th/Nb is used as a proxy to depict the contribution of slab components (Note: Ba/Nb would be better, but Ba is probably modified by alteration; Elliott, 2003; Pearce et al., 2005, 2006). The contribution of slab components become greater towards south (*Fig. 6h*), which is consistent with mantle wedge flow towards south and with northward subduction of the Neo-Tethyan seafloor.

We can thus conclude that the Mesozoic magmatism in the southern Lhasa Terrane probably resulted from northward subduction of the Neo-Tethyan seafloor, which supports the first model mentioned above, for which we illustrate in *Fig. 8*. We infer,

therefore, that the Neo-Tethyan seafloor subduction beneath the Lhasa Terrane must have commenced as early as middle Triassic (Wang et al., 2016), and the Neo-Tethyan ocean basin may have begun to develop in the Carboniferous – early Permian (e.g. Dewey et al., 1988).

5.4 Implications for Jurassic crustal accretion on the southern margin of the Lhasa Terrane

Juvenile crustal accretion on the southern margin of the Lhasa Terrane in the Mesozoic is most consistent with the magmatism caused by the northward subduction of the Neo-Tethys seafloor as we elaborate above. The juvenile crust so produced has been invoked as potential protoliths for the Miocene ore-bearing adakitic magmatism (cf. Hou et al., 2004, 2015; Zhu et al., 2011b, 2013). An important issue therefore concerns how the juvenile crust was accreted on the southern margin of the Lhasa Terrane. The SG and ZT volcanic associations with juvenile components are best explained as directly resulting from slab-dehydration induced mantle wedge melting, melt extraction and differentiation (Kang et al., 2014; Wang et al., 2016; Zhang et al., 2014; Zhu et al., 2009). However, we emphasize that juvenile crustal rocks represented by the YF volcanic rocks were most likely produced in the back-arc setting. Actually, the back-arc extension provides a possible scenario for crustal accretion, which is similar to that of the circum-Pacific orogens (Collins and Richards, 2008; Kemp et al., 2009). This back-arc accretionary model (cf. Collins, 2002; Collins and Richards, 2008) requires a tripartite association that comprises (1) inboard S-type granites, (2)

outboard magmatic arc and (3) intervening back-arc basin with sediments. Coincidentally, the late Triassic – early Jurassic magmatism in the southern Lhasa Terrane seems to meet these requirements: (1) the inboard S-type granites are represented by peraluminous granites (205-189Ma) in the central part of the Lhasa Terrane, which are best interpreted as being formed in collisional orogens (Liao et al., 2003; Liu et al., 2006; Zhang et al., 2007); (2) the volcanic sequences in the Sangri Group (194-137Ma; Kang et al., 2014; Zhu et al., 2009) and Zedong terrane (160-155Ma; Zhang et al., 2014) are possible candidates for the outboard magmatic arc; (3) the bimodal volcanic suite in the Yeba Formation (192-174Ma) we study here is most consistent with being formed in response to the initiation of the intervening back-arc basin. We, therefore, propose a late Triassic – early Jurassic crustal accretion model for the southern Lhasa Terrane, which is preliminary and will need refining and improving with more detailed petrological, geochemical and geochronological studies becoming available on magmatic rocks within the Gangdese batholiths.

A late Triassic, sediment-filled graben might have developed in the central part of the Lhasa Terrane as reflected by retrograde metamorphism in the Nyainqentanglha Group (230-213Ma; cf. Dong et al., 2011; Li et al., 2009; Zhu et al., 2013). This short-lived back-arc basin might have received and buried sediments to mid-crustal depths (cf. Collins, 2002; Collins and Richards, 2008). A new episode of back-arc rifting was re-established outboard resulting from trench retreat of the subducting Neo-Tethyan seafloor, which is identified by volcanic suite of the Yeba Formation.

During the trench retreat, hot basaltic magmas underplated the localities under extension, facilitating silicic magma generation as recorded by the S-type granites sporadically outcropped in the central part of the Lhasa Terrane (Liao et al., 2003; Liu et al., 2006; Zhang et al., 2007). With ongoing extension accompanied by crustal thinning, silicic magmas gradually lost their S-type character (cf. Collins and Richard, 2008) as reflected by the silicic volcanic rocks of the Yeba Formation with features of transitional I-S type granite. Cycles of tectonic activity like this can induce crustal accretion on the southern margin of the Lhasa Terrane.

6 Conclusions

1. The volcanic sequence in the Yeba Formation is compositionally bimodal, dominated by basalt/basaltic andesite and dacite/rhyolite dated at 174-183Ma.
2. The arc-like Yeba basalts are consistent with 5-15% partial melting of a mantle source metasomatized/enriched by subduction components. The silicic rocks, with characteristics of transitional I-S type granites, are most consistent with their parental magmas resulting from anatexis of a mixed source of juvenile amphibole-rich lower crust with reworked crustal materials compositionally similar to metagreywackes.
3. The YF volcanism is best interpreted as genetically associated with the initiation of a back-arc basin, which further implies northward subduction of the

Neo-Tethyan seafloor beneath the Lhasa Terrane. Importantly, such a back-arc rifting setting may be likely a site for net crustal accretion because the juvenile crustal materials can be preserved in the collisional orogen such as in the southern Lhasa Terrane.

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Figure Captions

Fig.1 Simplified geological map showing (a) tectonic outline of the Tibetan Plateau; (b) main tectonic units and distribution of Mesozoic volcanism in the Lhasa Terrane; (c) sample locations; (d) outcrop of the Yeba volcanic rocks; (e-g) photomicrographs of representative samples. Pl = plagioclase; Q = quartz

Fig.2 Rock classification diagrams of (a) Nb/Y vs. Zr/TiO_2 (after Winchester and Floyd, 1976) and (b) Co vs. Th (after Hastie et al., 2007). Literature data of YF volcanic rocks are from Zhu et al., (2008). IAT, island arc tholeiitic; CA, calc-alkaline; H-K/SHO, high potassic/shoshonitic; BA/A, basaltic andesite/andesite; D/R, dacite/rhyolite.

Fig.3 Chondrite-normalized REE patterns for (a) the YB basaltic rocks and (b) the YB silicic rocks; (c) N-MORB-normalized multi-element diagram for the YB basaltic rocks; (d) Primitive mantle-normalized multi-element diagram for the YB silicic rocks. Data of the Gangdese arc basalts are from Kang et al., (2014) and Wang et al., (2016); Okinawa Trough BABBs are from Shinjo et al., (2000). The values of chondrite, primitive mantle and N-MORB are from Sun and McDonough, 1989. BABB, back-arc basin basalt

Fig.4 Zircon U-Pb dating results, CL images and concordia plots for the YB volcanic rocks. Scale bar = 50 μm.

Fig. 5 (a) Zircon Hf and (b) whole rock Sr-Nd isotopic compositions for the YB volcanic rocks. Fields for the Gangdese batholiths and the Lhasa basement are from Ji et al., (2009) and Wu et al., (2010), respectively.

Fig.6 (a-e) Chemical variation diagrams for the YB volcanic rocks; (f) Partial melting model for the YB basalts; Spatial variations of (g) Nb/Ta and (h) Th/Nb show polarities of subduction and mantle convection. Crustal and GLOSS compositions are from Rudnick and Gao (2003) and Plank and Langmuir (1998); N-MORB value is from Sun and McDonough., (1989). See text for details.

Fig.7 Chemical compositions for the YB silicic volcanic rocks. Outlined regions and solid lines are from Altherr and Siebel., (2002), and Patiño Douce (1999) that depict compositions of partial melts obtained by dehydration experiments for various whole rock compositions. MB, meta-basalts; MA, meta-andesites; MGW, metagreywackes; MP, metapelites; MMP, meta-mafic apelites; MFP, meta-felsic apelites.

Fig.8 Schematic model to show the Jurassic volcanism at the southern margin of

the Lhasa Terrane. Modified after Liu et al., (2015). CC, continental crust; SCLM, subcontinental lithospheric mantle

Table captions

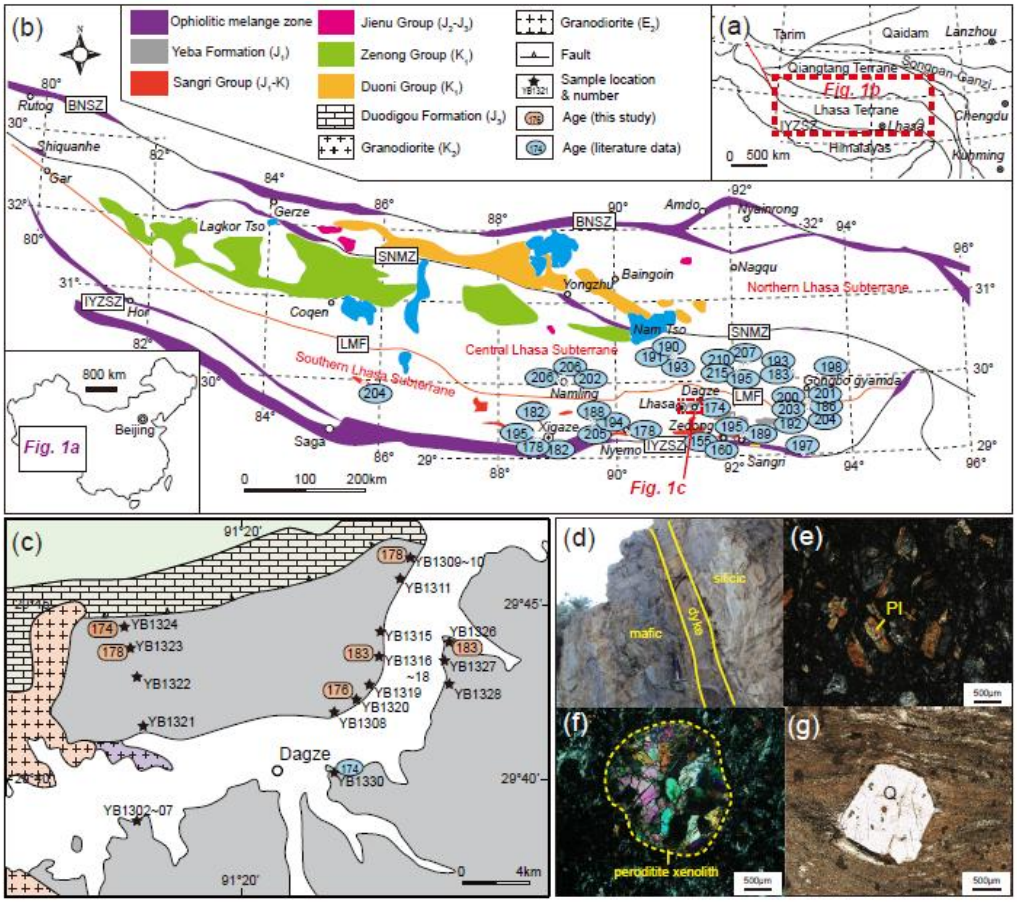
Table 1 Whole-rock geochemical data for the volcanics of Yeba Formation

- LOI, loss of ignition; $Mg^{\#}$ = molar 100×MgO/(MgO+FeO_T), FeO_T=0.9×$_TFe_2O_3$, ASI = molar Al_2O_3/(Na_2O+K_2O+CaO)

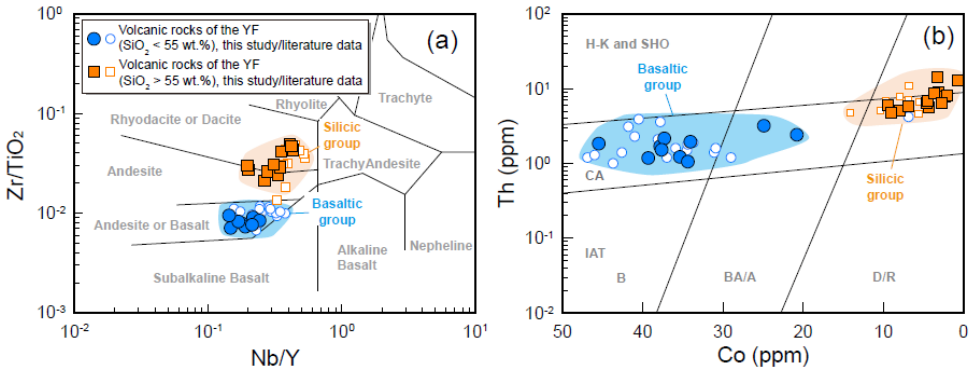
- $\varepsilon_{Nd}(t) = [(^{143}Nd/^{144}Nd)s / (^{143}Nd/^{144}Nd)_{CHUR} - 1] \times 10000$, $T_{DM} = \ln[(^{143}Nd/^{144}Nd)s - (^{143}Nd/^{144}Nd)_{DM}] / [(^{143}Sm/^{144}Nd)s - (^{147}Sm/^{144}Nd)_{DM}] / \lambda$.

$(^{143}Nd/^{144}Nd)_{CHUR}$ = 0.512638, $(^{147}Sm/^{144}Nd)_{CHUR}$ = 0.1967, $(^{143}Nd/^{144}Nd)_{DM}$ = 0.51315, $(^{147}Sm/^{144}Nd)_{DM}$ = 0.2136 and t = 174 Ma

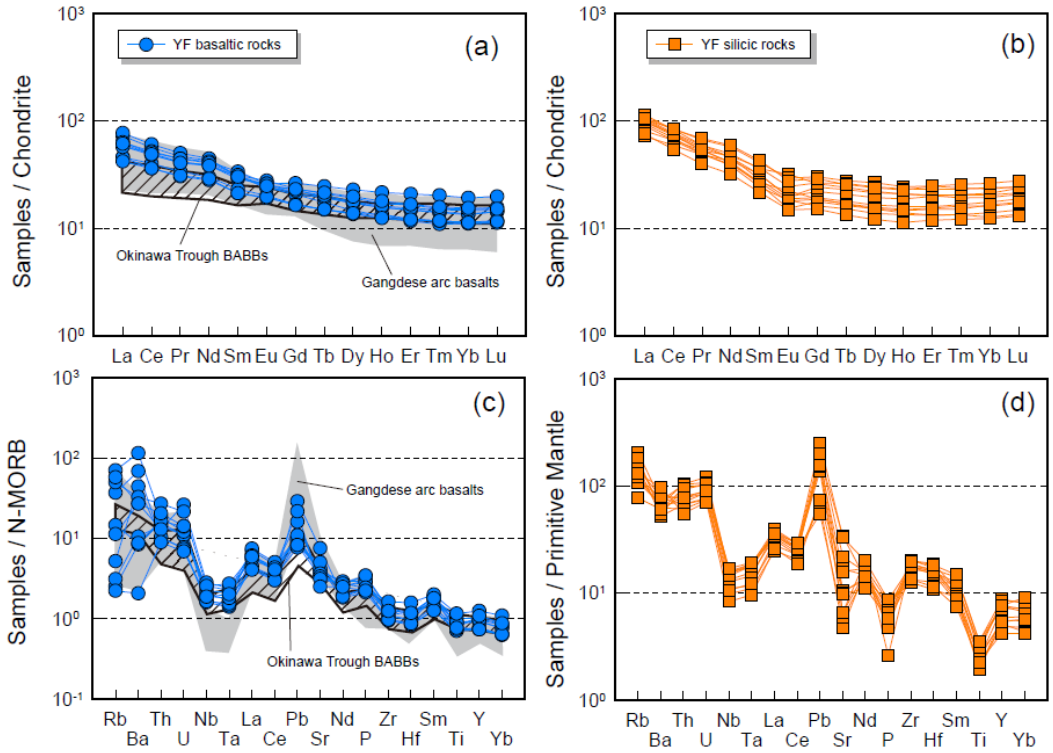
Wei et al. Fig. 1 W160 mm - H142 mm (2-column fitting image)



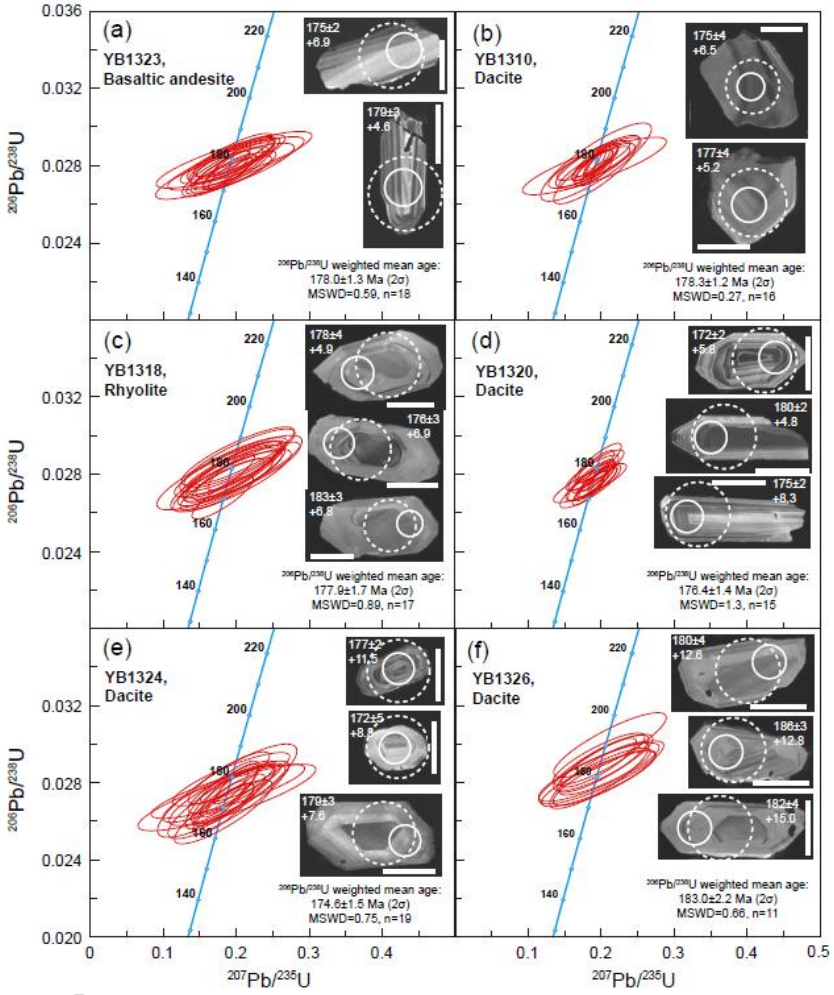
Wei et al. Fig. 2 W156 mm - H59 mm (2-column fitting image)



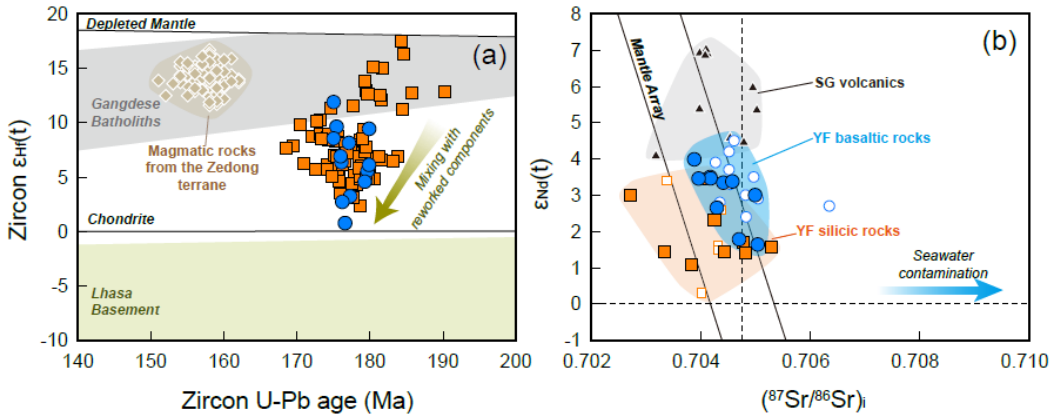
Wei et al. Fig. 3 W156 mm - H112 mm (2-column fitting image)



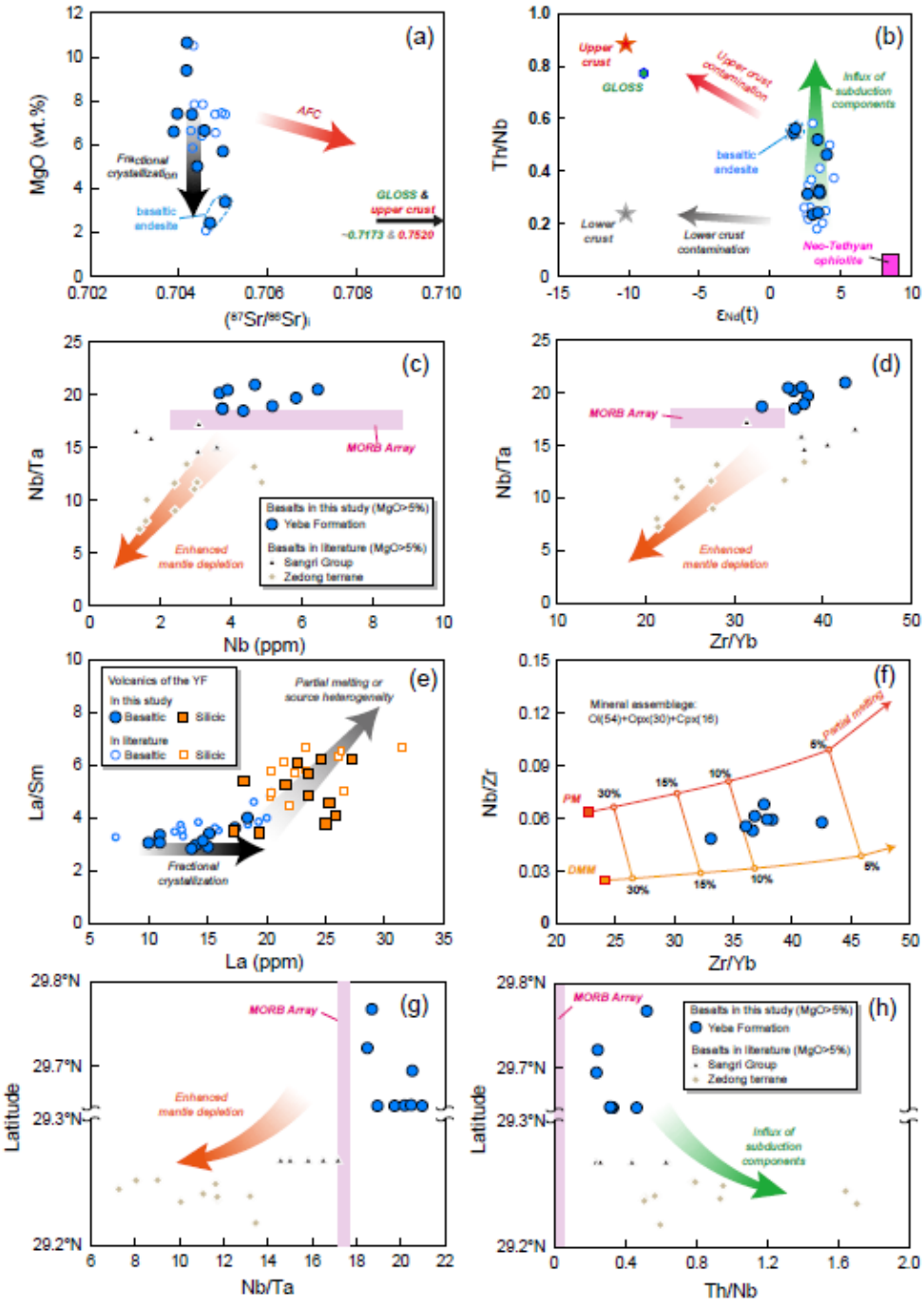
Wei et al. Fig. 4 W145 mm - H173 mm (2-column fitting image)



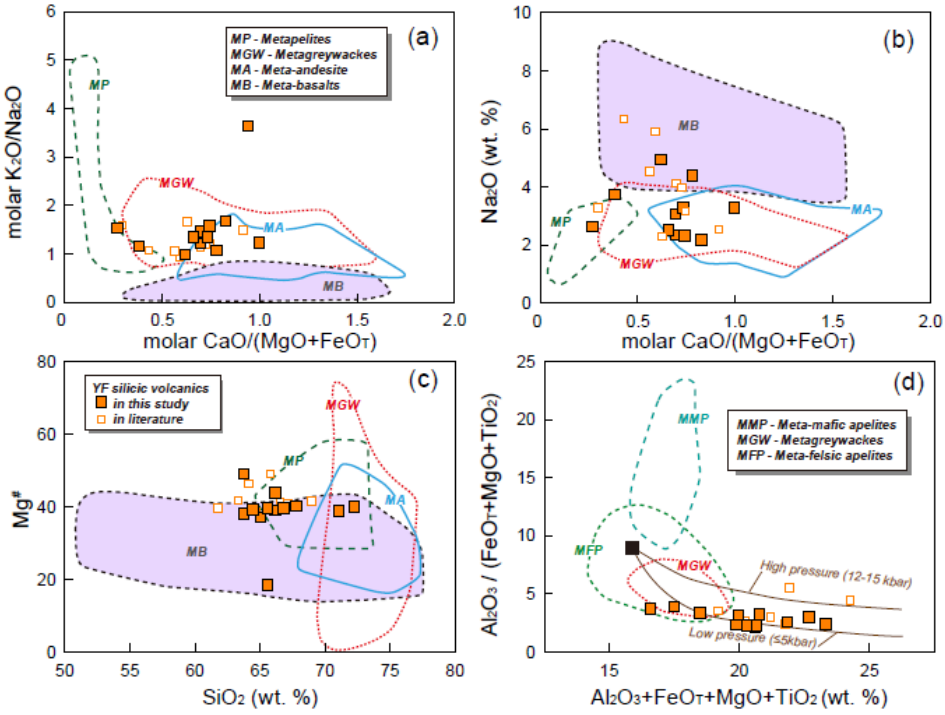
Wei et al. Fig. 5 W154 mm - H61 mm (2-column fitting image)



Wei et al. Fig. 6 W168 mm - H237 mm (2-column fitting image)



Wei et al. Fig. 7 W156 mm - H117 mm (2-column fitting image)



Wei et al. Fig. 8 W160 mm - H100 mm (2-column fitting image)

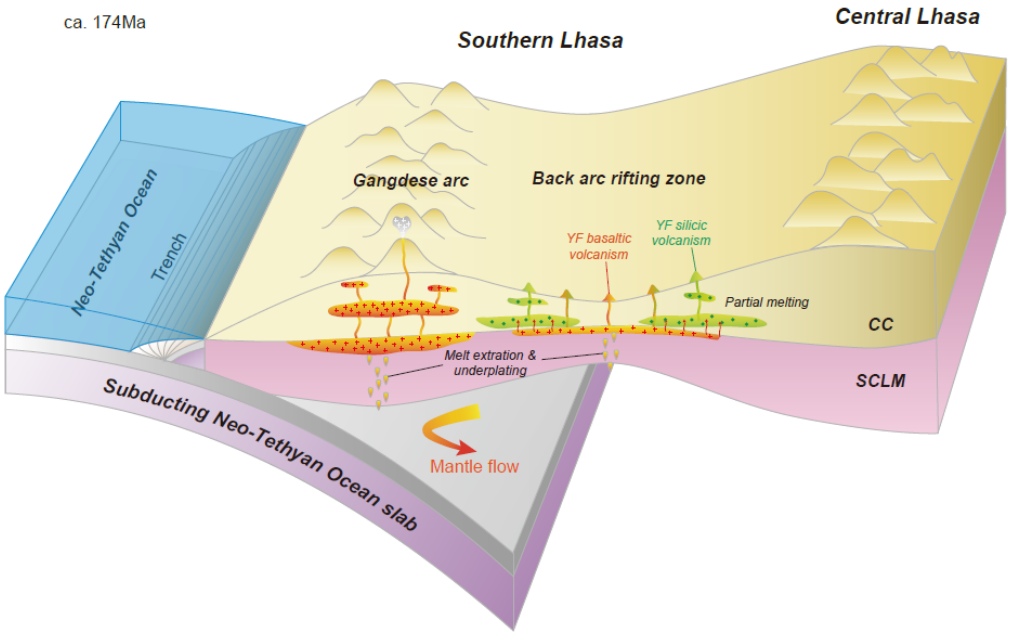


Table 1 Whole-rock geochemical data for the volcanics of Yeba Formation

Sample	YB1302	YB1303	YB1304	YB1306	YB1307	YB1311	YB1321	YB1322	YB1319	YB1323	YB1308
Rock type	Basalt	Basalt	Basalt	Basalt	Basalt	Basalt	Basalt	Basalt	B.Andesite	B.Andesite	Dacite
Latitude(N)	29°39.1'	29°39.1'	29°39.1'	29°39.1'	29°39.2'	29°46.1'	29°41.7'	29°43.3'	29°43.2'	29°44.2'	29°41.9'
Longitude(E)	91°16.5'	91°16.5'	91°16.5'	91°16.7'	91°16.8'	91°25.8'	91°16.9'	91°16.9'	91°24.9'	91°16.7'	91°23.0'
Age (Ma, 2σ)										178.0±1.3	
SiO_2	41.46	40.83	47.87	48.03	48.82	47.93	50.58	47.56	52.08	52.49	65.02
TiO_2	1.26	1.14	0.90	0.88	0.96	1.09	1.13	0.93	1.44	0.96	0.72
Al_2O_3	16.86	17.89	17.14	16.58	16.83	18.06	18.00	16.09	14.89	15.30	13.91
TFe_2O_3	13.35	11.20	10.15	10.26	10.01	10.89	8.99	9.45	10.92	10.60	5.00
MnO	0.15	0.19	0.15	0.16	0.21	0.18	0.28	0.18	0.19	0.30	0.09
MgO	9.39	10.65	6.60	7.42	7.38	5.02	5.70	6.65	3.40	2.45	1.47
CaO	7.48	7.59	8.16	9.34	5.96	7.35	2.96	10.58	5.13	7.45	3.86
Na_2O	3.30	2.90	2.84	2.24	5.05	4.35	6.57	1.47	4.46	2.17	3.06
K_2O	0.10	0.12	2.15	1.01	0.24	0.30	0.26	0.07	0.84	1.04	2.36
P_2O_5	0.39	0.34	0.34	0.24	0.27	0.33	0.40	0.27	0.28	0.26	0.17
LOI	5.52	6.17	3.63	3.52	3.62	3.89	4.20	6.10	5.76	6.14	3.36
Total	99.24	99.03	99.93	99.68	99.34	99.38	99.06	99.32	99.40	99.13	99.02
ASI	0.88	0.95	0.78	0.76	0.87	0.87	1.09	0.74	0.85	0.84	0.95
$Mg^{\#}$	58	65	56	59	59	48	56	58	38	31	37
Sc	40.8	36.2	32.2	33.8	34.6	31.0	35.0	28.3	31.9	28.1	15.9
V	190	242	277	259	272	330	226	240	336	217	91.8
Cr	211	184	162	212	223	44.9	84.8	169	2.51	14.2	9.04
Co	45.5	37.8	37.3	39.3	35.4	34.1	37.6	34.4	24.9	20.8	9.43
Ni	81.1	70.0	72.3	91.4	67.7	29.2	43.0	85.7	10.7	8.23	5.08
Zn	109.1	88.2	76.2	76.7	72.1	82.3	99.1	75.6	141	111	50.4
Ga	11.3	13.2	16.3	15.7	12.9	18.0	15.6	15.9	17.6	16.0	14.7
Rb	1.39	1.72	38.7	20.3	2.86	8.07	6.23	1.23	27.4	31.6	73.5
Sr	347	307	351	369	445	323	297	666	278	222	357
Y	27.9	25.3	21.5	20.3	20.5	25.4	26.5	20.3	34.6	29.9	35.0
Zr	98.4	86.8	80.8	69.2	70.2	77.3	94.7	70.8	119	90.7	194
Nb	5.83	5.16	4.67	3.67	3.90	3.75	6.45	4.34	5.85	4.31	7.04
Ba	52.0	65.4	711	424	274	54.0	168	12.8	203	167	470
La	17.3	15.1	18.4	10.9	10.9	14.0	13.6	9.98	15.0	14.6	17.2
Ce	34.8	30.6	37.3	22.3	24.7	29.5	31.3	22.2	32.0	30.1	36.6
Pr	4.46	3.97	4.78	2.91	3.23	3.93	4.24	2.95	4.24	3.88	4.62
Nd	20.1	18.1	20.9	13.2	14.1	18.3	19.1	13.6	19.6	17.9	20.0
Sm	4.77	4.43	4.58	3.26	3.57	4.69	4.81	3.26	5.17	4.65	4.90
Eu	1.61	1.51	1.41	1.11	1.17	1.49	1.37	1.14	1.51	1.44	1.14
Gd	4.83	4.34	4.26	3.45	3.43	4.71	4.63	3.38	5.42	4.78	4.99
Tb	0.79	0.72	0.66	0.56	0.56	0.73	0.76	0.56	0.92	0.81	0.86
Dy	4.87	4.37	3.73	3.45	3.51	4.38	4.50	3.51	5.82	5.02	5.46
Ho	0.97	0.89	0.75	0.71	0.77	0.89	0.91	0.71	1.23	1.02	1.19
Er	2.73	2.45	2.01	1.94	2.03	2.48	2.56	2.00	3.47	2.78	3.38
Tm	0.39	0.36	0.30	0.28	0.30	0.36	0.38	0.30	0.52	0.40	0.52
Yb	2.57	2.29	1.90	1.89	1.95	2.34	2.52	1.92	3.28	2.63	3.50
Lu	0.40	0.36	0.30	0.29	0.30	0.36	0.37	0.29	0.50	0.39	0.54
Hf	2.39	2.22	1.95	1.70	1.82	2.01	2.34	1.76	3.20	2.37	5.09
Ta	0.30	0.27	0.22	0.18	0.19	0.20	0.31	0.23	0.35	0.26	0.46
Pb	2.58	3.06	2.82	2.83	2.26	3.22	2.44	8.54	4.76	6.40	5.32
Th	1.85	1.70	2.16	1.18	1.23	1.95	1.52	1.06	3.19	2.42	6.02
U	0.50	0.51	0.56	0.36	0.34	0.65	1.21	0.32	0.99	0.66	1.83
$^{87}Sr/^{86}Sr$	0.704204	0.704224	0.704714	0.704377	0.704346	0.704604	0.705152	0.704592	0.705788	0.705780	
$(^{87}Sr/^{86}Sr)_i$	0.704174	0.704183	0.703883	0.703963	0.704298	0.704416	0.704994	0.704579	0.705043	0.704703	
$^{143}Nd/^{144}Nd$	0.512755	0.512758	0.512767	0.512760	0.512723	0.512760	0.512740	0.512751	0.512678	0.512682	
$(^{143}Nd/^{144}Nd)_i$	0.512588	0.512586	0.512613	0.512586	0.512545	0.512580	0.512563	0.512582	0.512493	0.512500	

$\varepsilon_{Nd}(t)$	3.50	3.46	3.99	3.46	2.65	3.35	3.00	3.38	1.64	1.78
T_{DM}(Ga)	0.86	0.91	0.72	0.93	1.08	1.01	1.02	0.89	1.32	1.25

Table 1 (continued)

Sample	YB1309	YB1310	YB1315	YB1316	YB1318	YB1320	YB1324	YB1326	YB1327	YB1328	YB1330
Rock type	Dacite	Dacite	Dacite	Dacite	Rhyolite	Dacite	Dacite	Dacite	Rhyolite	Dacite	Dacite
Latitude(N)	29°46.4'	29°46.2'	29°44.1'	29°43.4'	29°43.3'	29°42.5'	29°44.3'	29°44.1'	29°44.1'	29°43.5'	29°40.3'
Longitude(E)	91°25.8'	91°25.8'	91°24.9'	91°24.8'	91°24.8'	91°24.3'	91°16.6'	91°26.3'	91°26.3'	91°26.2'	91°23.1'
Age (Ma, 2σ)		178.3±1.2			177.9±1.7	176.4±1.4	174.6±1.5	183.3±2.2			
SiO_2	63.75	66.17	67.77	66.81	71.02	66.18	65.59	64.45	72.22	65.55	63.73
TiO_2	0.59	0.72	0.77	0.47	0.43	0.54	0.62	0.53	0.42	0.78	0.49
Al_2O_3	16.38	15.69	14.37	14.22	13.91	15.21	13.88	15.82	13.07	17.00	14.11
TFe_2O_3	5.25	4.45	4.46	3.10	2.58	3.27	4.39	3.61	2.50	4.84	4.12
MnO	0.07	0.12	0.17	0.11	0.09	0.11	0.13	0.08	0.08	0.06	0.09
MgO	1.61	1.42	1.49	1.01	0.82	1.27	1.43	1.16	0.83	0.54	1.97
CaO	4.14	1.95	3.44	3.57	2.30	3.36	4.77	3.04	1.80	1.12	4.19
Na_2O	2.35	3.74	2.52	3.30	4.39	2.17	0.56	3.28	4.95	2.64	2.33
K_2O	2.76	2.45	2.48	2.60	2.19	3.41	2.80	3.20	1.71	3.47	3.25
P_2O_5	0.18	0.18	0.19	0.13	0.11	0.12	0.15	0.14	0.11	0.06	0.13
LOI	2.61	2.39	2.45	4.34	2.46	4.64	6.10	4.39	2.44	2.95	5.42
Total	99.67	99.29	100.11	99.65	100.29	100.28	100.42	99.69	100.13	99.00	99.82
ASI	1.14	1.27	1.10	0.96	1.01	1.14	1.10	1.10	0.98	1.68	0.94
$Mg^{\#}$	38	39	40	39	39	44	39	39	40	18	49
Sc	11.0	11.9	15.7	6.56	6.35	9.96	15.5	7.94	6.04	14.6	12.2
V	89.6	35.0	38.3	40.0	33.0	49.9	82.1	46.3	28.1	46.3	85.4
Cr	7.42	4.24	4.71	1.62	1.34	1.34	8.31	1.80	1.17	6.30	18.6
Co	7.96	4.35	4.61	4.44	3.10	3.71	6.83	2.29	2.06	2.77	9.06
Ni	4.51	2.84	3.09	1.72	1.59	1.74	4.96	1.96	1.49	3.33	8.49
Zn	57.7	87.1	82.2	54.7	49.9	59.2	74.8	67.8	45.9	47.0	48.4
Ga	15.4	17.4	16.2	13.8	13.4	16.2	14.3	15.6	11.5	19.9	13.3
Rb	106	96.4	67.7	80.5	69.2	131	76.7	116	49.2	91.6	83.1
Sr	724	338	700	222	462	139	101	126	214	140	209
Y	23.5	34.4	40.1	22.2	27.0	32.4	38.8	25.6	24.6	35.1	19.3
Zr	142	151	203	198	212	225	187	230	199	224	151
Nb	7.87	9.14	11.1	9.56	11.0	11.4	7.66	11.1	10.4	12.2	5.97
Ba	440	367	570	428	468	510	537	387	424	554	683
La	21.6	25.3	25.8	22.6	27.3	23.5	19.3	23.6	24.6	25.0	18.0
Ce	40.5	47.7	51.8	42.6	51.2	46.8	40.7	45.0	46.3	51.6	33.5
Pr	4.68	5.66	6.36	4.72	5.69	5.39	5.06	5.07	5.14	6.57	3.79
Nd	18.5	23.2	26.9	18.0	21.4	21.5	22.2	19.4	19.2	27.8	15.0
Sm	4.09	5.51	6.30	3.72	4.37	4.84	5.62	4.13	3.96	6.60	3.32
Eu	1.30	1.86	1.71	1.06	1.10	1.15	1.43	1.04	0.89	1.62	0.87
Gd	3.87	5.36	6.24	3.34	3.97	4.53	5.47	3.84	3.67	5.96	3.11
Tb	0.63	0.91	1.04	0.57	0.67	0.78	0.98	0.64	0.61	0.96	0.50
Dy	3.87	5.63	6.79	3.51	4.19	4.92	6.21	4.01	3.87	6.07	3.08
Ho	0.79	1.18	1.39	0.76	0.91	1.07	1.33	0.85	0.83	1.29	0.64
Er	2.20	3.34	4.10	2.21	2.64	3.29	3.82	2.55	2.50	4.10	1.97
Tm	0.34	0.49	0.61	0.34	0.43	0.53	0.59	0.41	0.39	0.66	0.31
Yb	2.17	3.31	3.97	2.38	2.99	3.62	4.09	2.73	2.83	4.49	2.10
Lu	0.35	0.53	0.63	0.38	0.48	0.60	0.61	0.44	0.44	0.71	0.33
Hf	3.35	3.92	5.16	4.63	5.06	5.85	4.86	5.32	4.93	5.66	3.55
Ta	0.48	0.60	0.72	0.64	0.77	0.77	0.50	0.74	0.74	0.79	0.40
Pb	13.4	11.6	17.9	13.6	14.2	9.90	4.71	3.91	10.9	11.2	5.21
Th	5.18	5.74	6.54	6.83	8.89	8.78	5.81	7.93	8.12	6.54	4.72
U	1.56	1.58	1.80	1.87	2.55	2.27	1.61	2.17	2.19	1.92	1.49

$^{87}Sr/^{86}Sr$	0.705167	0.706355	0.705489	0.707101	0.706403	0.710763	0.708326	0.710100	0.706519
$(^{87}Sr/^{86}Sr)_i$	0.704093	0.704252	0.704776	0.704425	0.705300	0.703844	0.702720	0.703337	0.704825
$^{143}Nd/^{144}Nd$	0.512742	0.512696	0.512662	0.512630	0.512634	0.512624	0.512741	0.512635	0.512629
$(^{143}Nd/^{144}Nd)_i$	0.512590	0.512533	0.512501	0.512488	0.512495	0.512469	0.512568	0.512489	0.512487
$\varepsilon_{Nd}(t)$	3.44	2.31	1.70	1.45	1.58	1.08	3.00	1.46	1.42
T_{DM}(Ga)	0.78	0.99	1.03	0.89	0.87	1.04	1.03	0.93	0.89

- LOI, loss of ignition; B. Andesite, Basaltic andesite; $Mg^{\#}$ = molar $100\times MgO/(MgO+FeO_T)$, $FeO_T=0.9\times {}_TFe_2O_3$, ASI = molar $Al_2O_3/(Na_2O+K_2O+CaO)$

- $\varepsilon_{Nd}(t) = [(^{143}Nd/^{144}Nd)s / (^{143}Nd/^{144}Nd)_{CHUR} - 1] \times 10000$, $T_{DM} = \ln[(^{143}Nd/^{144}Nd)s - (^{143}Nd/^{144}Nd)_{DM}] / [(^{143}Sm/^{144}Nd)s - (^{147}Sm/^{144}Nd)_{DM}] / \lambda$ (DePaolo, 1988). $(^{143}Nd/^{144}Nd)_{CHUR} = 0.512638$, $(^{147}Sm/^{144}Nd)_{CHUR} = 0.1967$, $(^{143}Nd/^{144}Nd)_{DM} = 0.51315$, $(^{147}Sm/^{144}Nd)_{DM} = 0.2136$ and t = 174 Ma

Highlights

1. New zircon U-Pb ages of 183-174 Ma are obtained for the Yeba Formation volcanics.
2. The Yeba volcanism was caused by the initiation of back-arc rifting.
3. Our results indicate northward Neo-Tethyan seafloor subduction beneath the Lhasa Terrane.
4. The back-arc setting is a likely site for juvenile crustal accretion in the southern Lhasa Terrane.